

Andy Sproles, ORNL

Microscopic models of nuclear structure at scale

Wouter Ryssens, G. Grams, M. Bender and S. Goriely

5th of June 2023

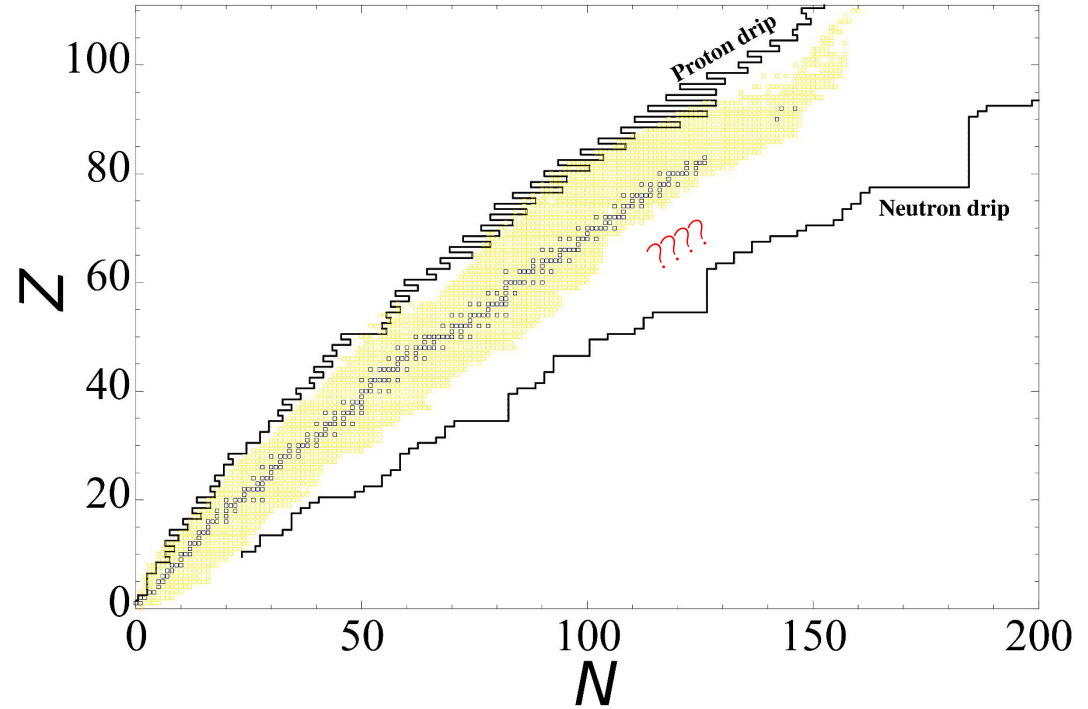


wryssens.com

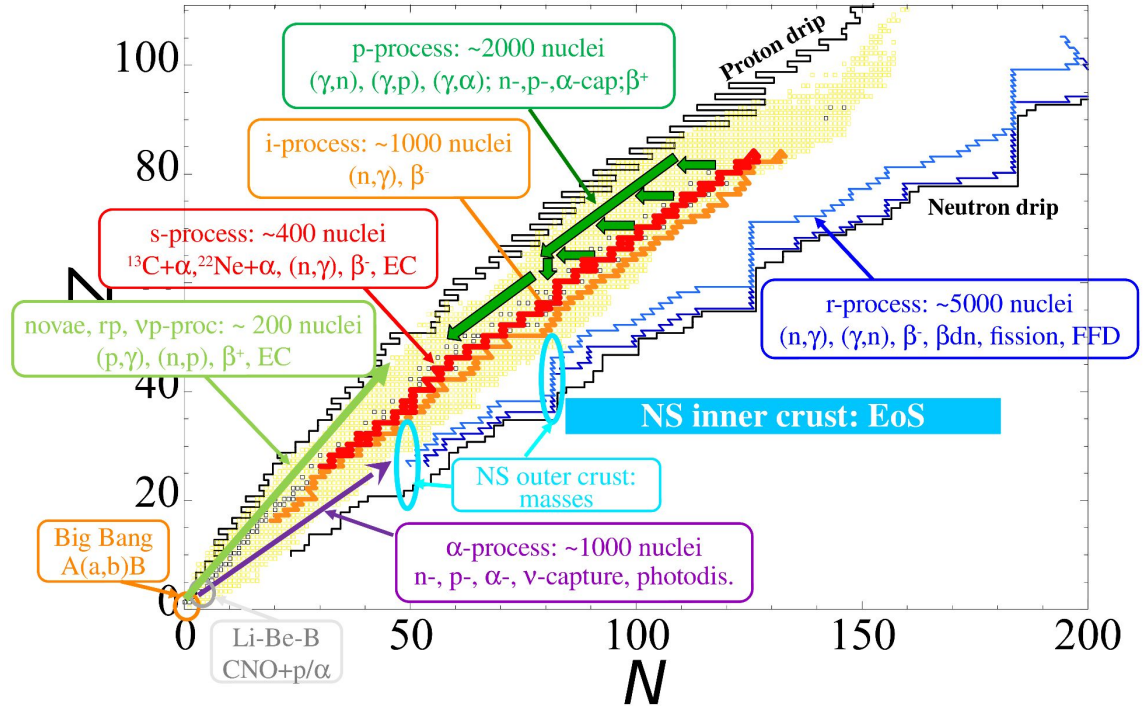
wryssens@ulb.be



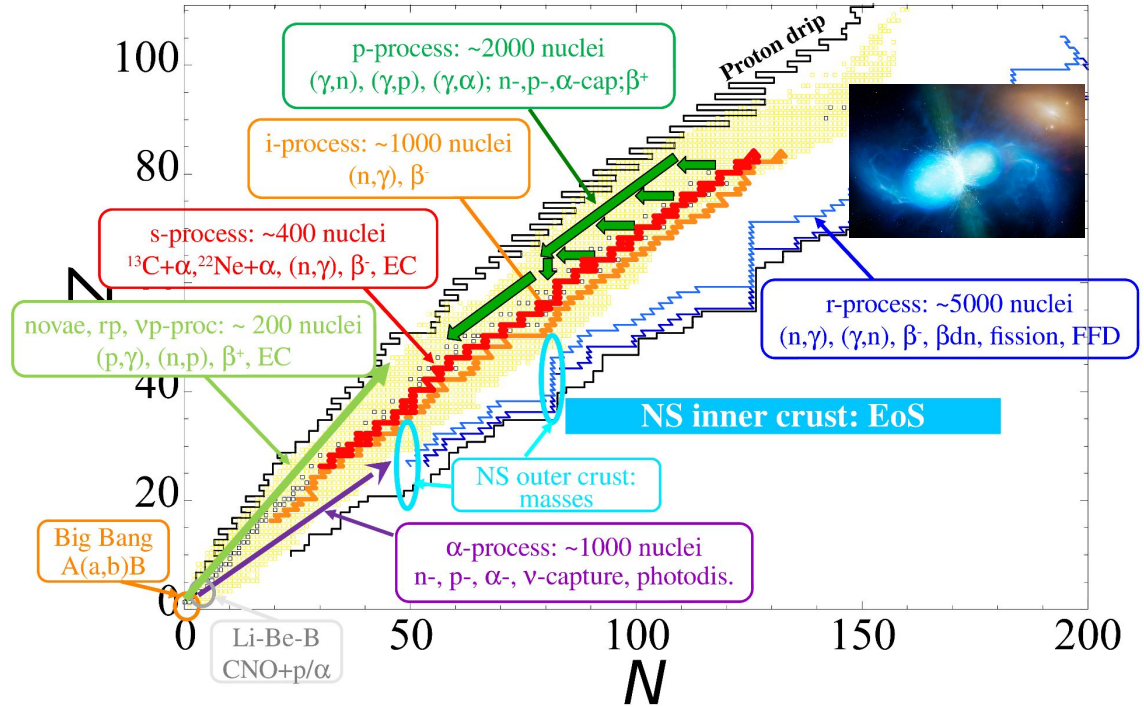
The nuclear chart...



The nuclear chart and the processes traversing it



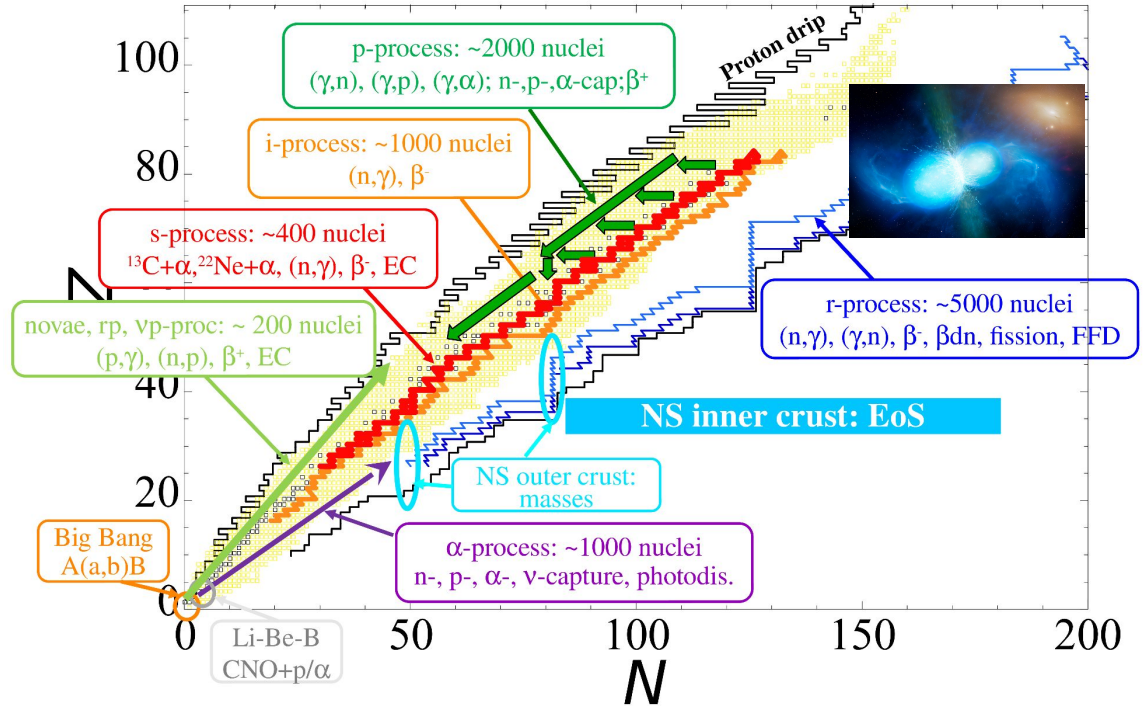
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Extrapolations in

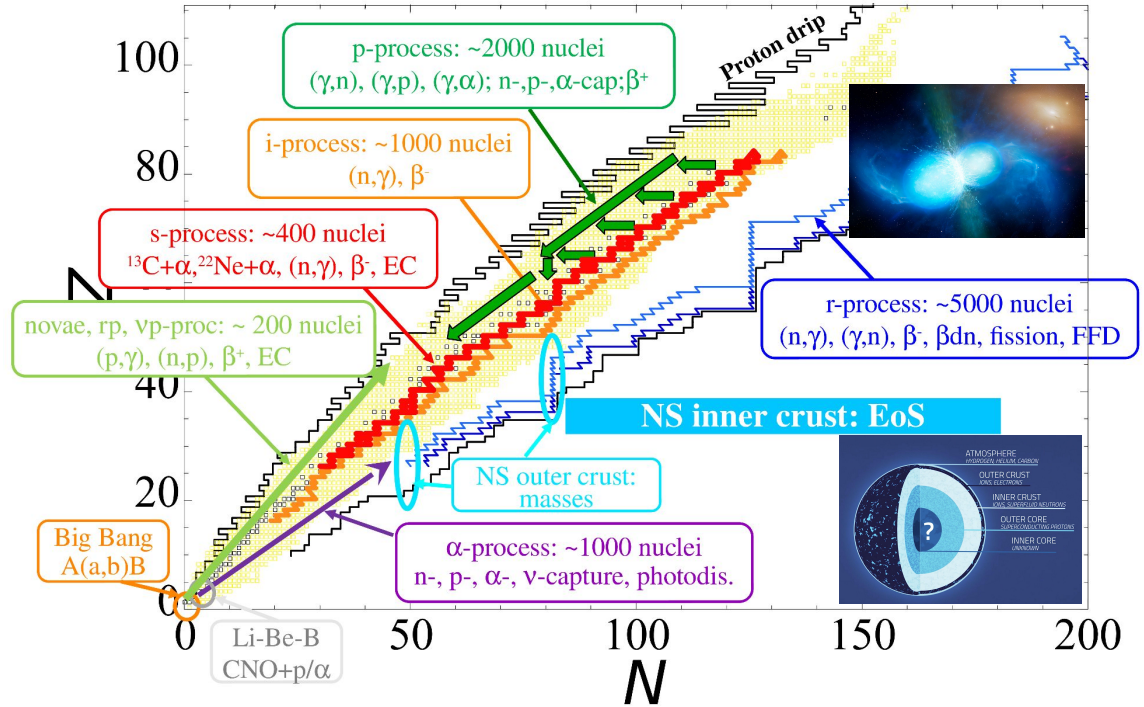
- nucleon number
- energy
- temperature
- density
-



The nuclear chart and the processes traversing it

Extrapolations in

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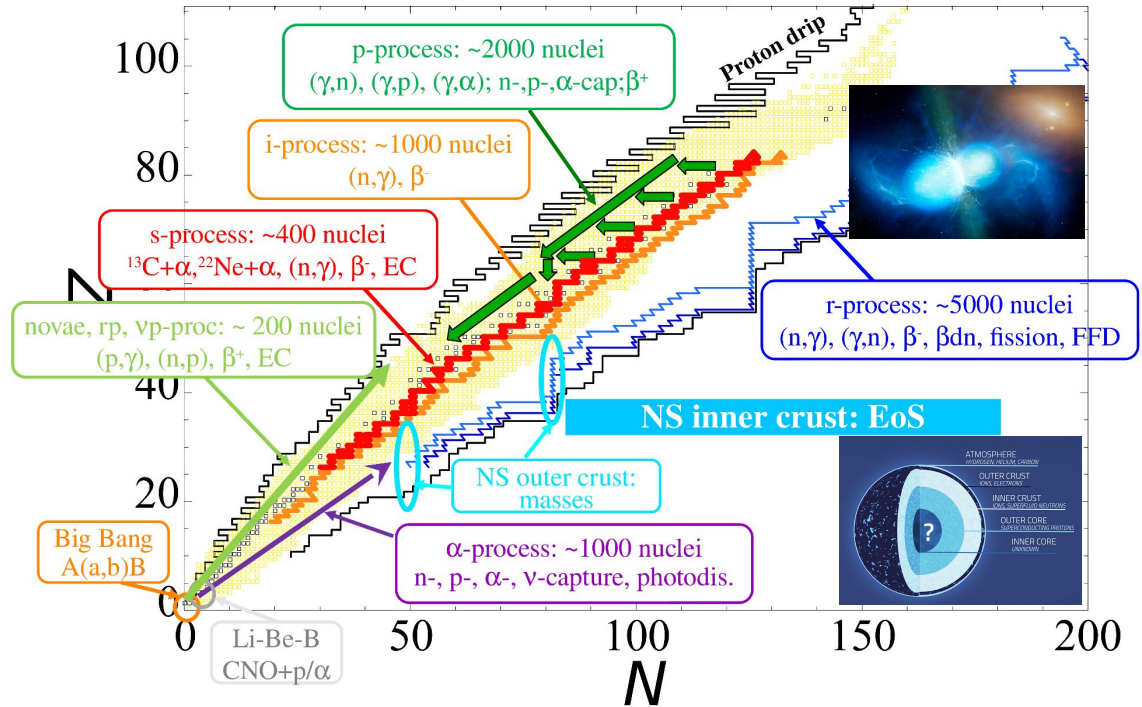
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and all of that for

- ~7000 nuclei
- many reactions



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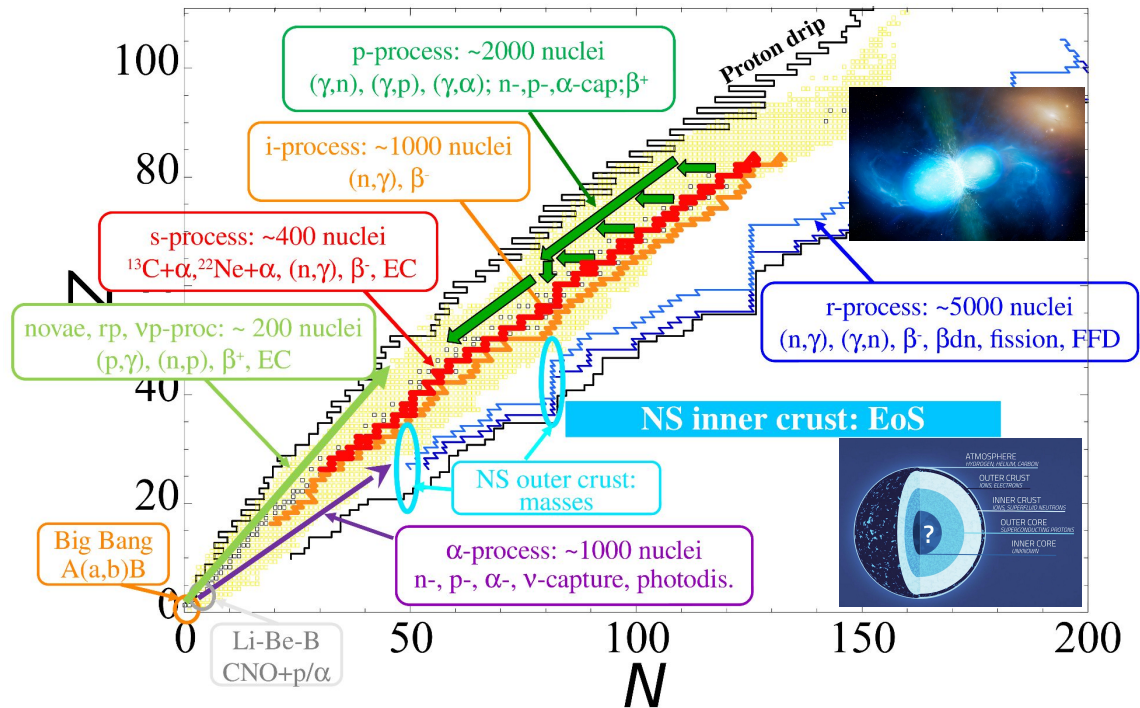
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and all of that for

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what we need is models that should be

1. predictive....
2. but also complete



Skyrme **E**nergy **D**ensity **F**unctionals (**EDFs**)

$$E \sim \int d^3r \left[C^\rho \rho(\mathbf{r})\rho(\mathbf{r}) + C^\tau \tau(\mathbf{r})\rho(\mathbf{r}) + \dots \right]$$



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Local densities and currents of a wavefunction

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Strong points

- wavefunctions with individual nucleons
- based on “in-medium” N-N interaction
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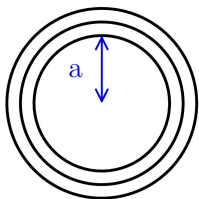
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How to move forward?

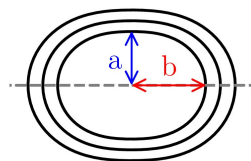
1. search for a “better” EDF form
2. include more experimental information
3. include more physics in the wavefunction

Large-scale models in 1-2 dimensions

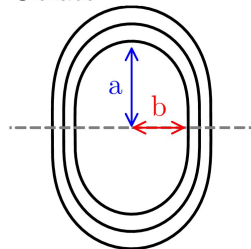
Spherical



Prolate



Oblate



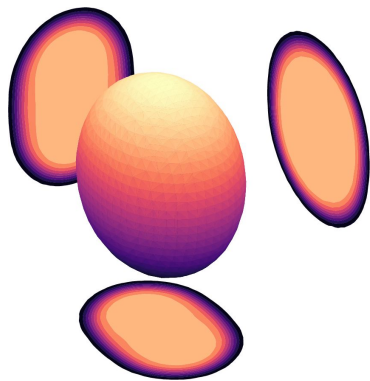
One DOF: β_{20}

Nuclear deformation

- larger variational space
- shape DOF characterized by multipole moments
- capture correlations at modest CPU cost
- intuitive interpretation

Large-scale models in 1-2-3 dimensions

β_{20}, β_{22} or β_2, γ



Symmetry breaking leads to deformation

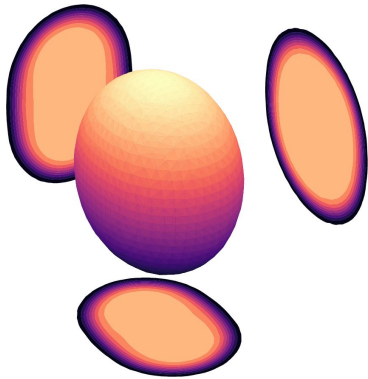
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More general configurations

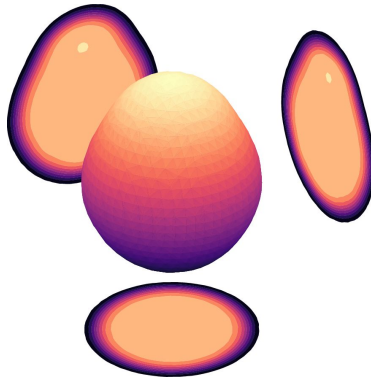
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β_{20}, β_{30}



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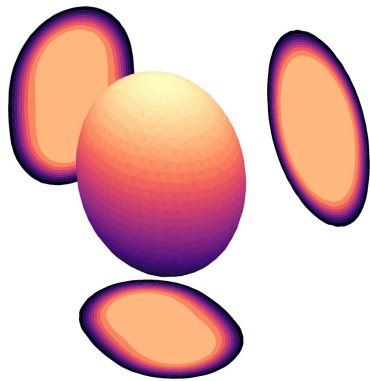
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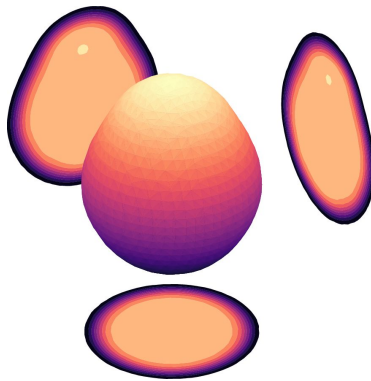
- triaxial shapes
- reflection asymmetry

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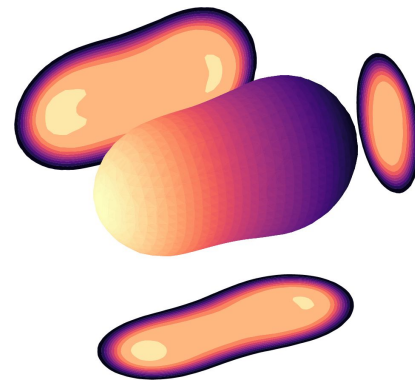
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β_{20}, β_{22} **and** β_{30}



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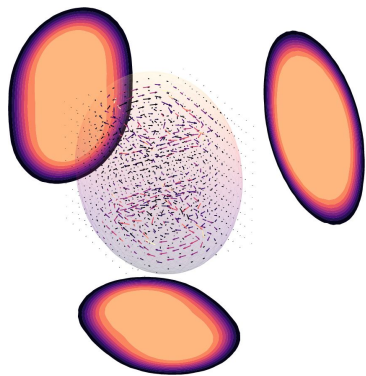
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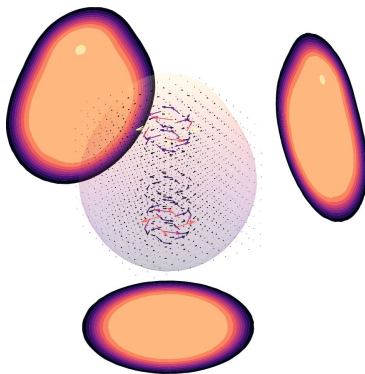
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- elongated shapes

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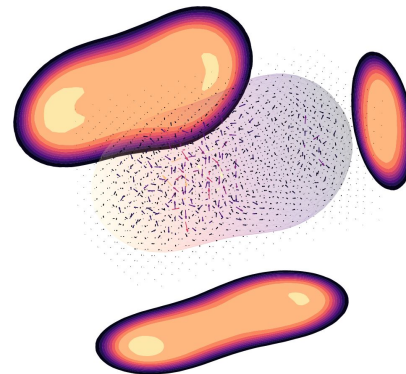
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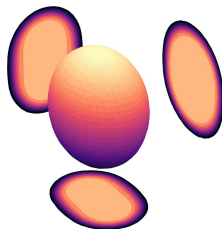
More general configurations

- triaxial shapes
- reflection asymmetry
- elongated shapes
- spin densities and currents

Brussels-Skyrme-on-a-Grid: BSkG

BSkG1 (2021)

- fitted to 2457 masses
- fitted to 884 charge radii
- includes triaxial deformation



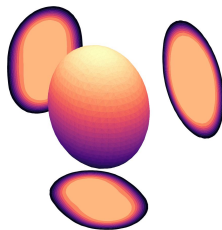
BSkG1: G. Scamps et al., EPJA **57**, 333 (2021).
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BSkG3: G. Grams et al., in preparation.

Rms σ	BSkG1	BSkG2	BSkG3
Masses [MeV]	0.741		
Radii [fm]	0.024		

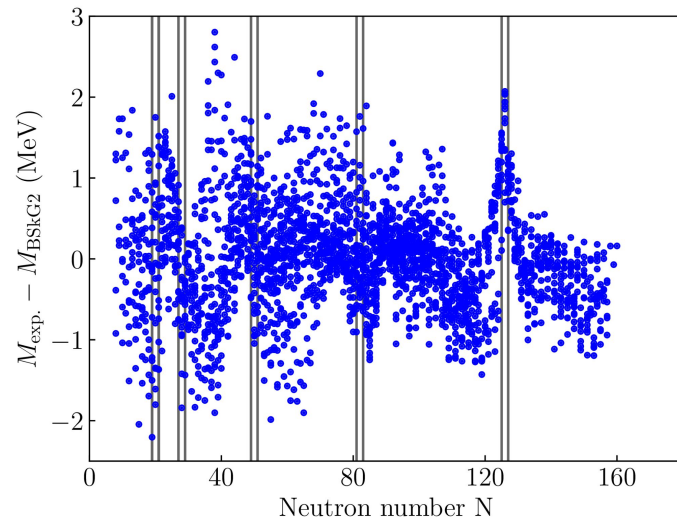
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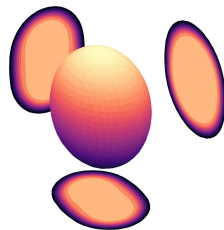
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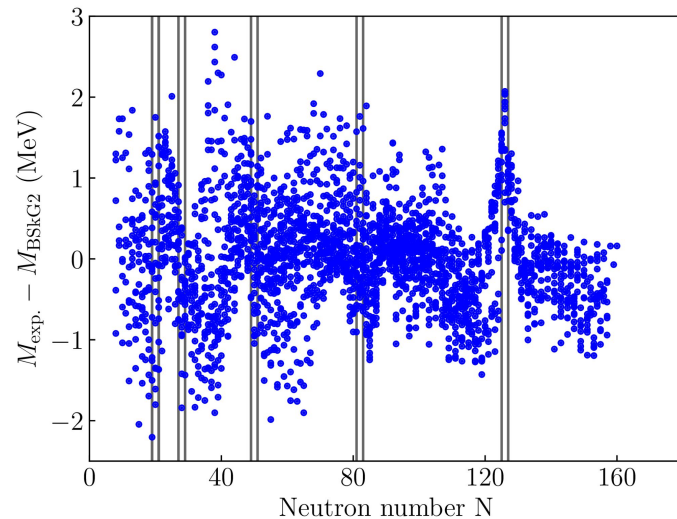
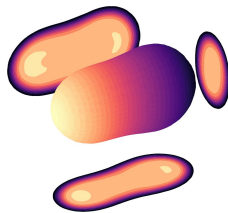
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BSkG2 (2022)

- fitted to 45 fission barriers
- includes spins, currents,...



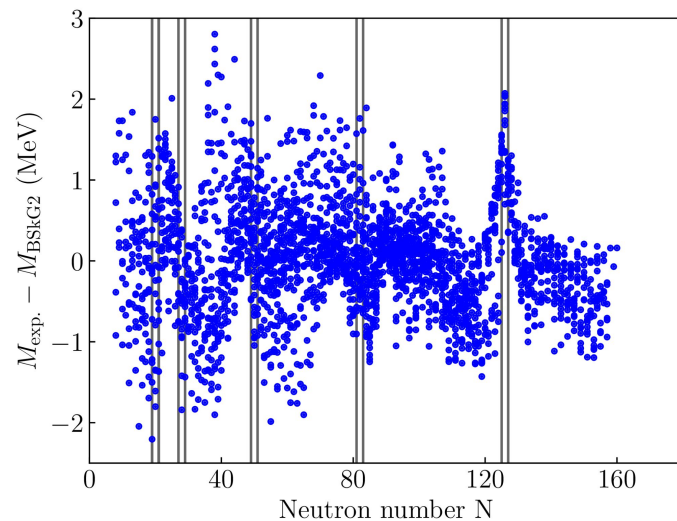
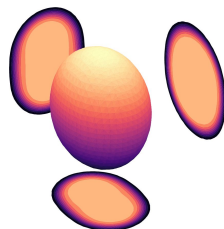
Rms σ	BSkG1	BSkG2	BSkG3
Masses [MeV]	0.741	0.678	
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Prim. barriers [MeV]	0.88	0.44	
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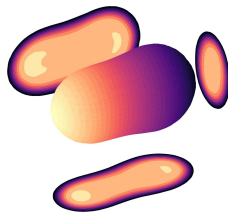
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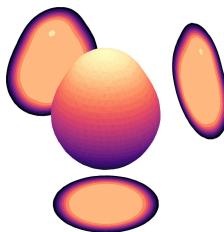
BSkG2 (2022)

- fitted to 45 fission barriers
- includes spins, currents,...



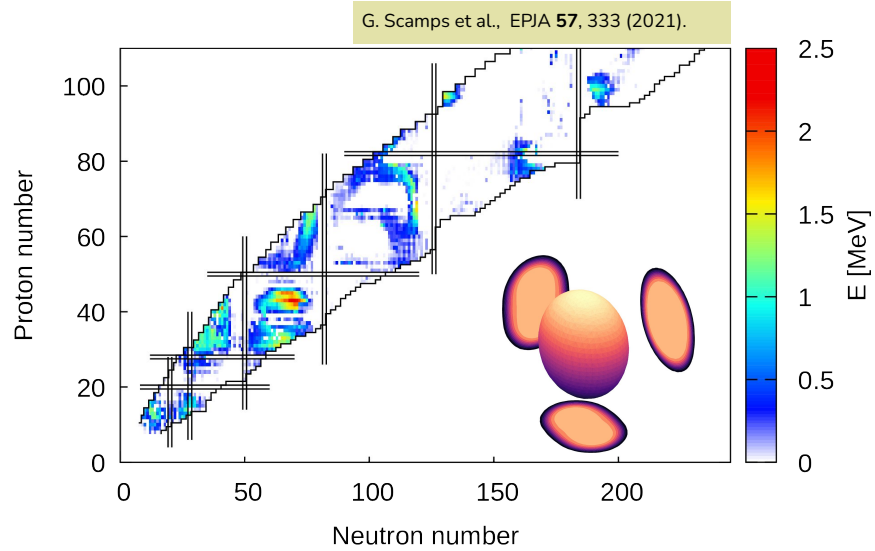
BSkG3 (2023)

- larger max. neutron star mass
- includes octupole deformation



Rms σ	BSkG1	BSkG2	BSkG3
Masses [MeV]	0.741	0.678	0.631
Radii [fm]	0.024	0.027	0.024
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Fission isomers [MeV]	1.0	0.49	0.34
Max. NS mass [M_{\odot}]	1.8	1.8	2.3

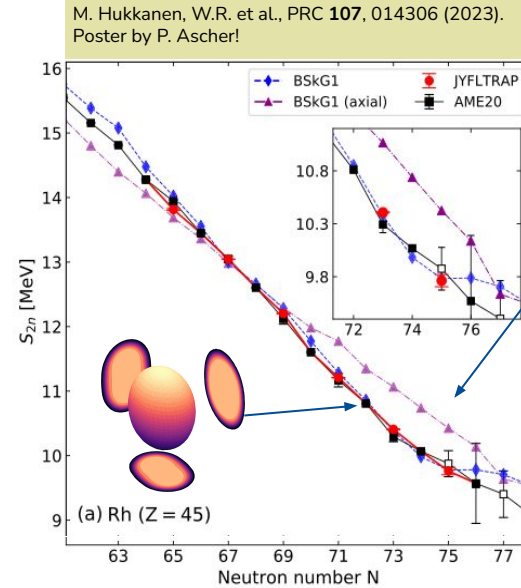
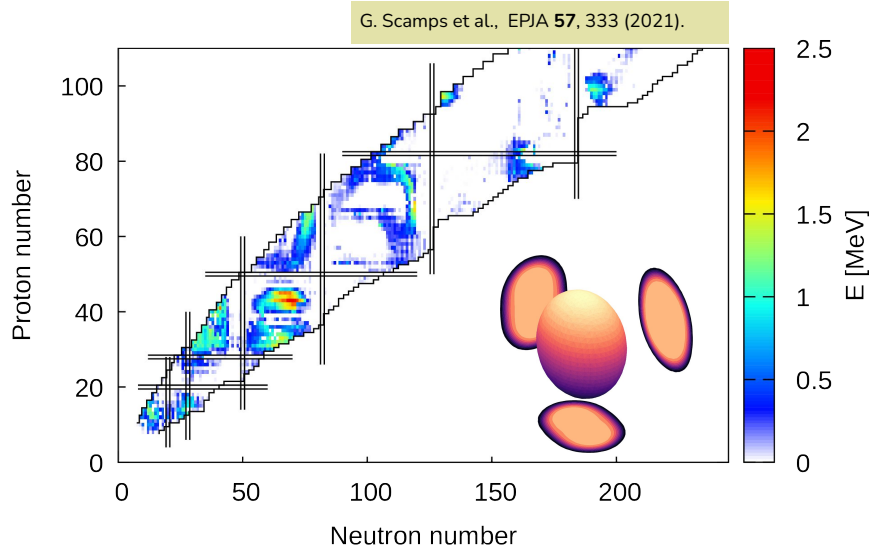
Masses



Triaxial deformation

- many nuclei are affected
- effects up to 2.5 MeV near $Z \sim 44$

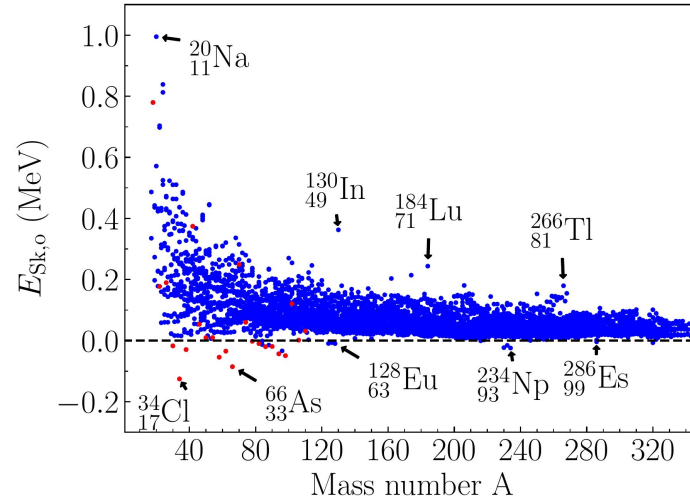
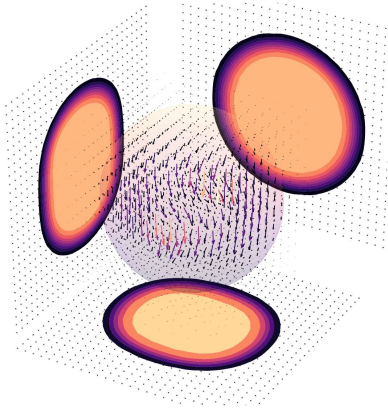
Masses



Triaxial deformation

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- effects up to 2.5 MeV near $Z \sim 44$
- does help reproduce trends, e.g. Rh

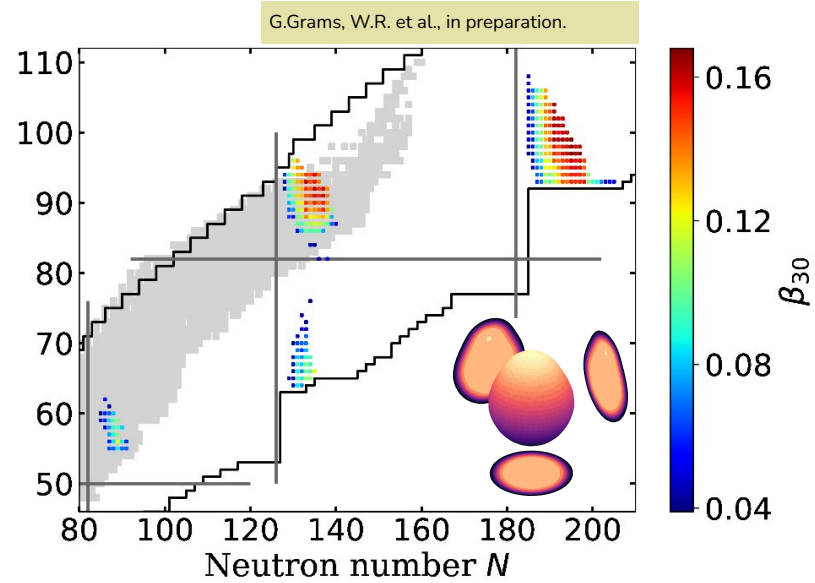
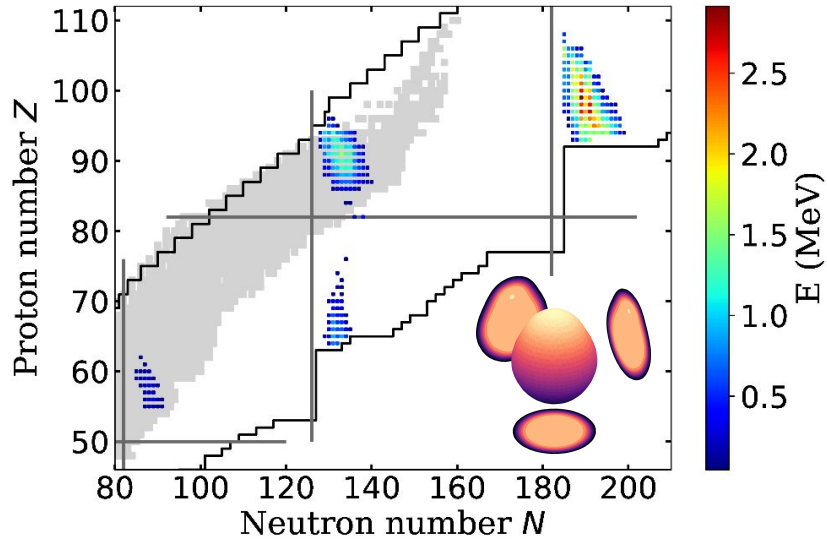
Masses



Time-odd terms

- small impact on the masses
- globally repulsive
- first time checked on this scale!
- first step towards other observables

Masses



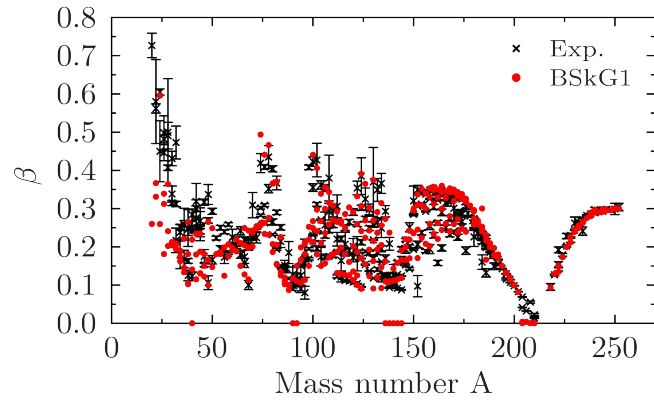
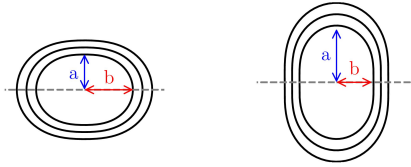
Reflection asymmetry

- small number of known nuclei affected
- Near $N=184$:
 - large effect up to 2.5 MeV
 - dripline modified
 - fission properties modified

Deformations

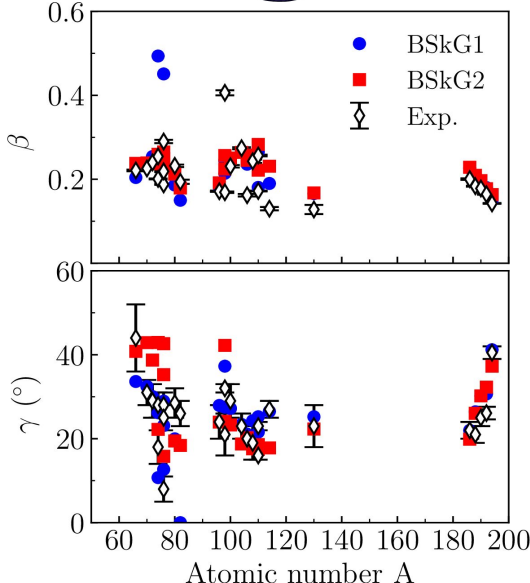
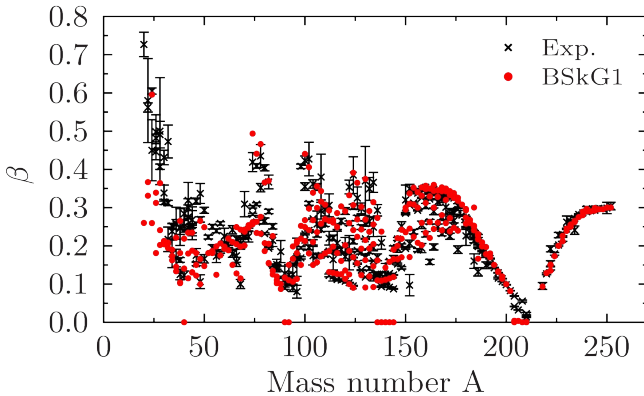
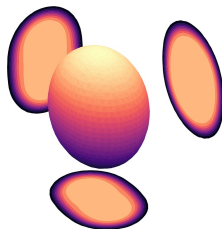
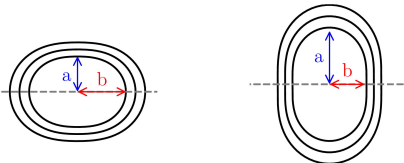
Deformations

“Ordinary” quadrupole deformation



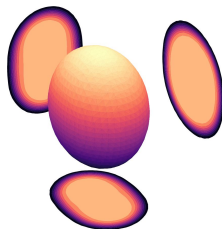
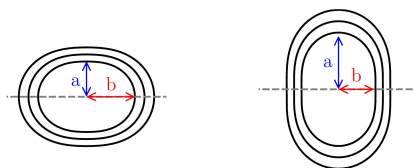
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“Ordinary” quadrupole deformation ... and triaxial deformation ...

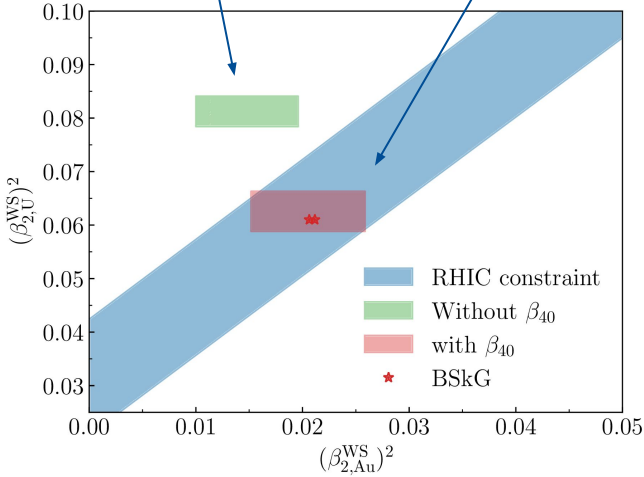
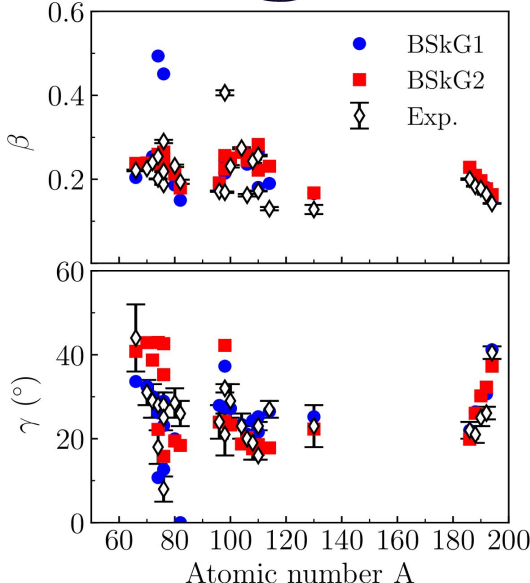
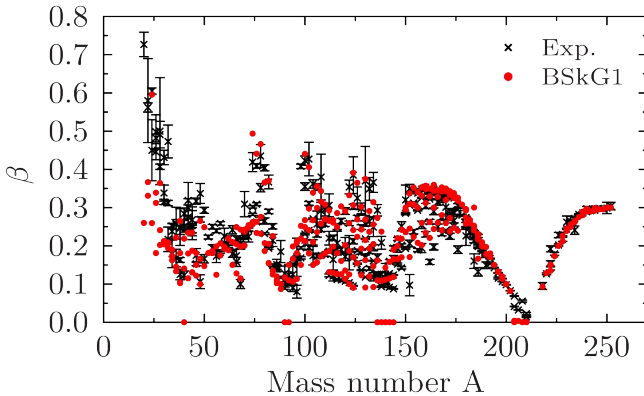


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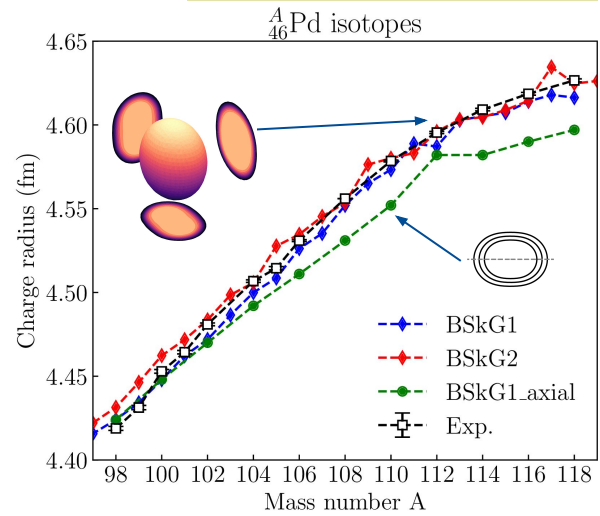


... and even hexadecapole!



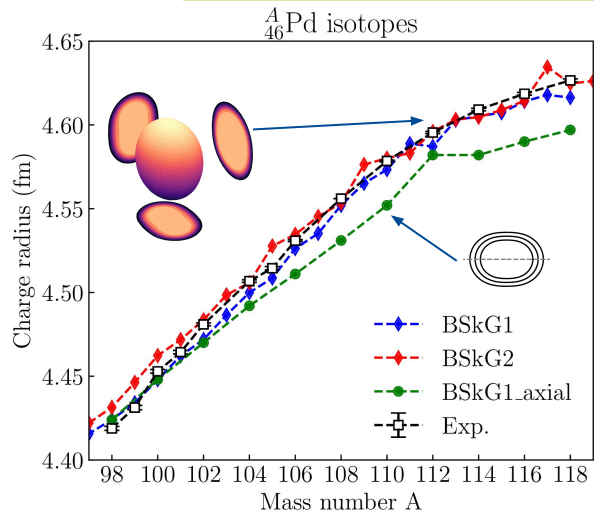
Radii

S. Geldhof, PRL **128**, 152501 (2022).
Talk by S. Geldhof on Thursday!

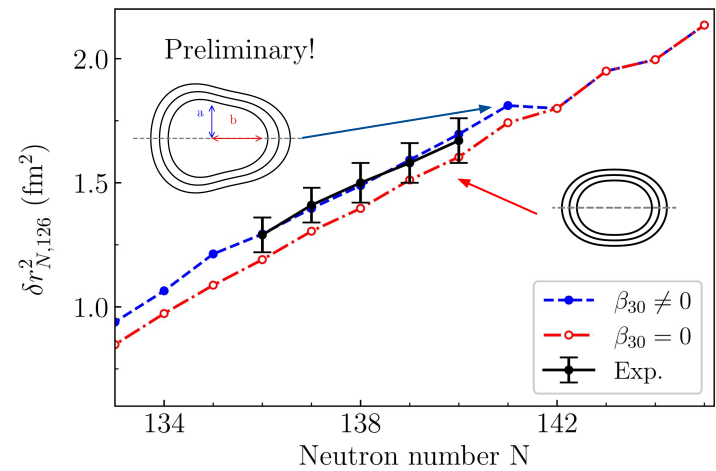


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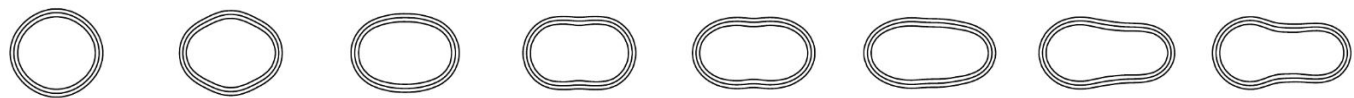
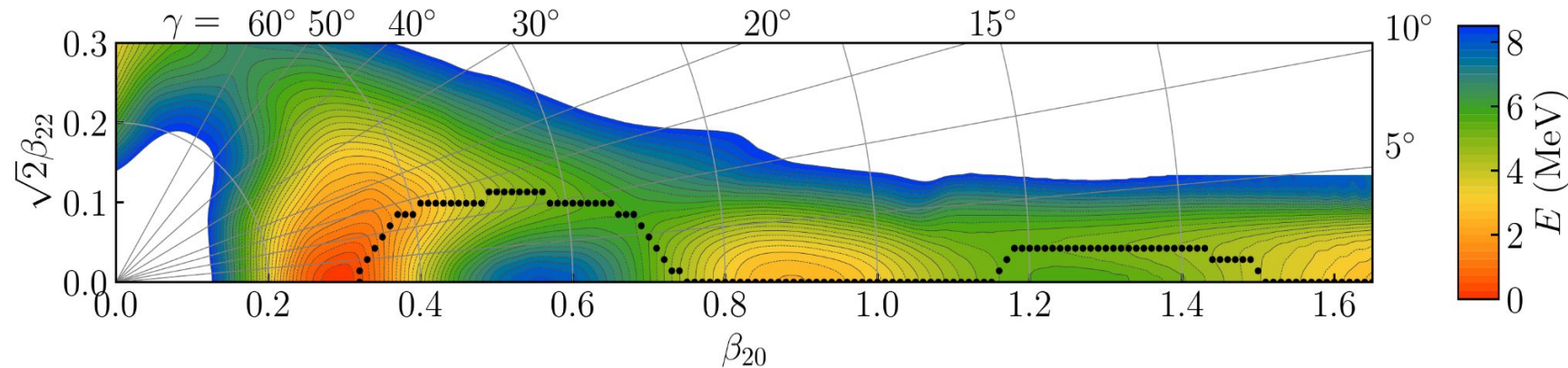
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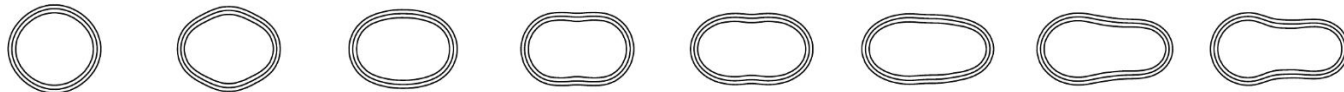
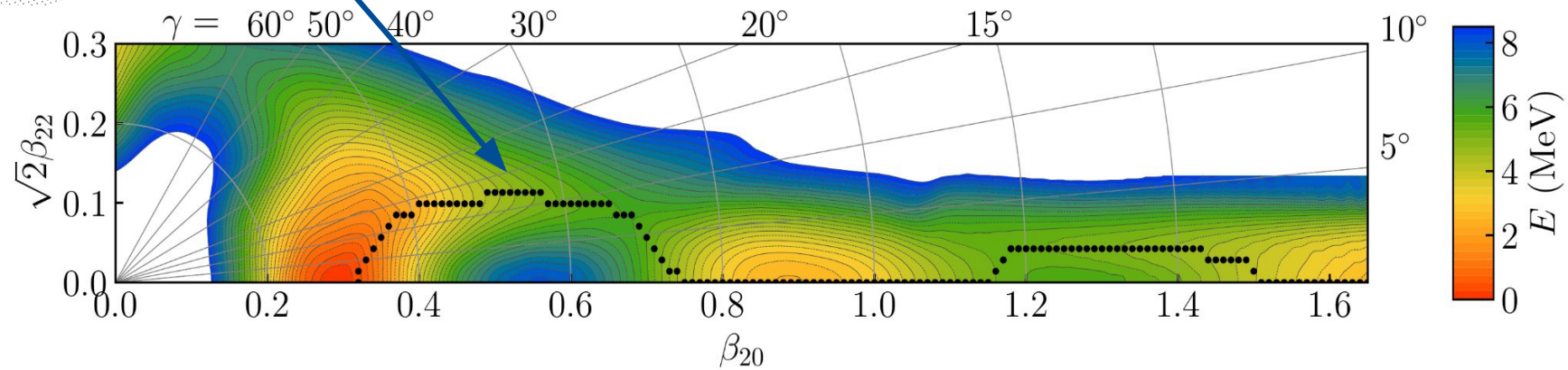
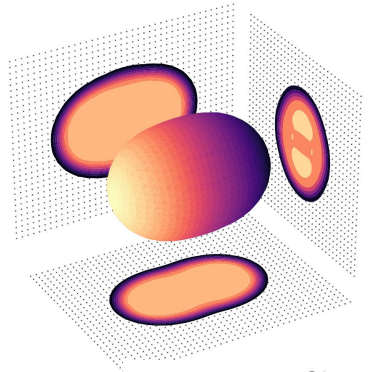
E. Verstraelen, PRC 100, 044321 (2019)



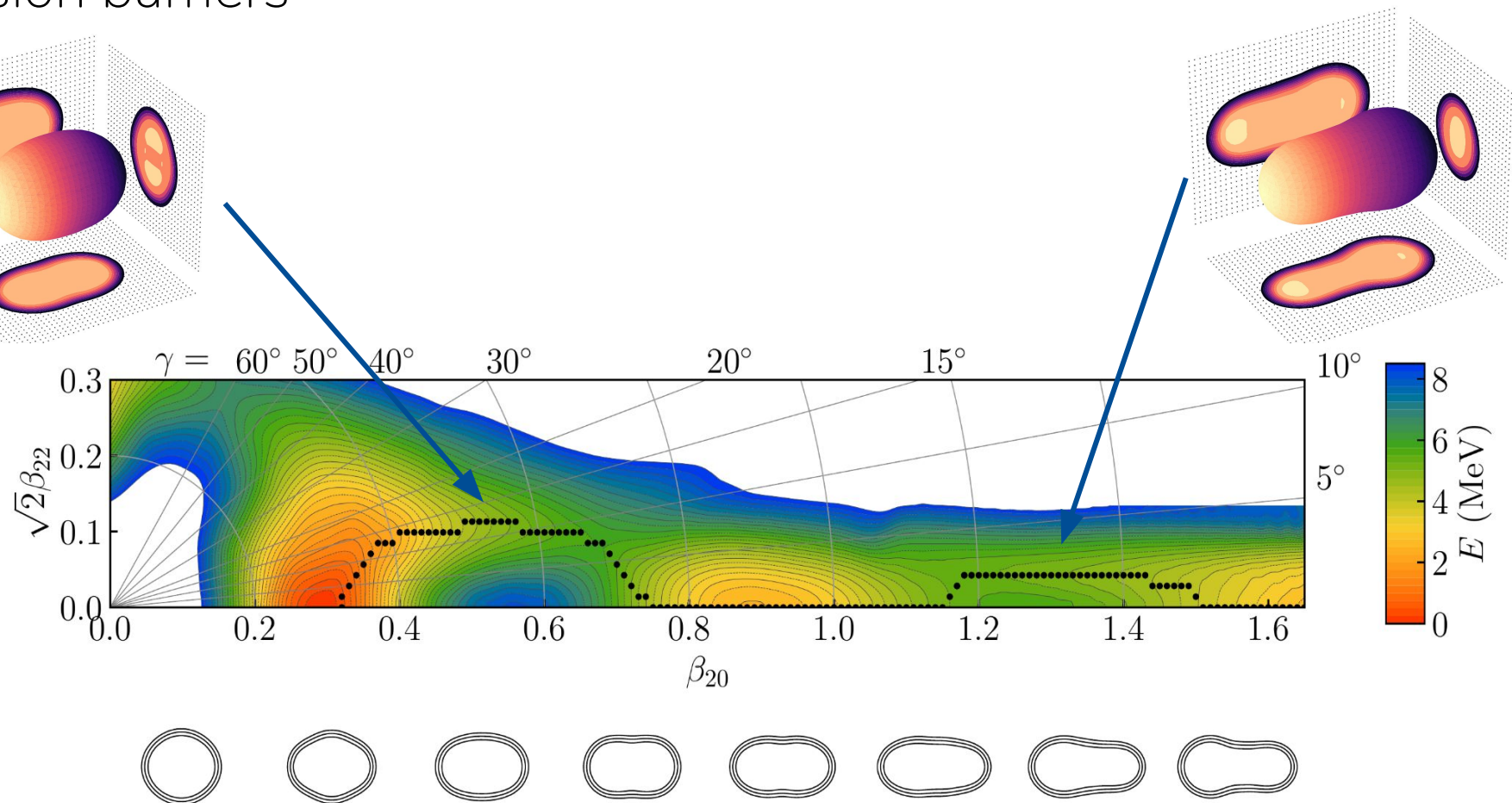
Fission barriers



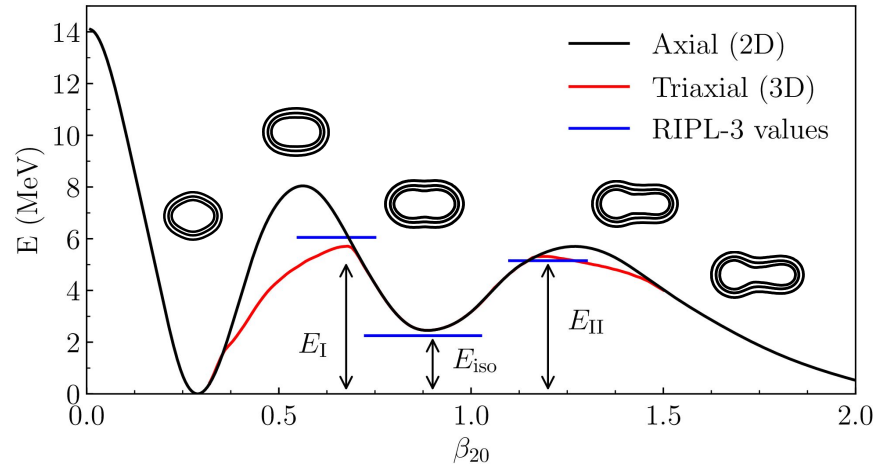
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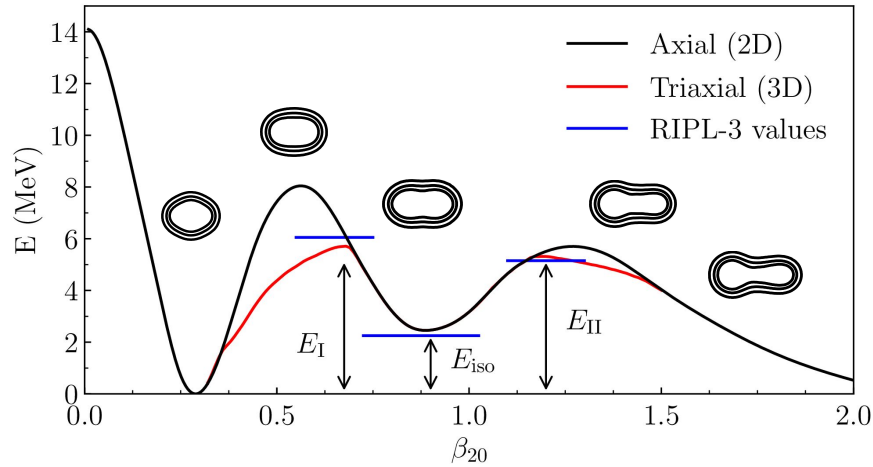
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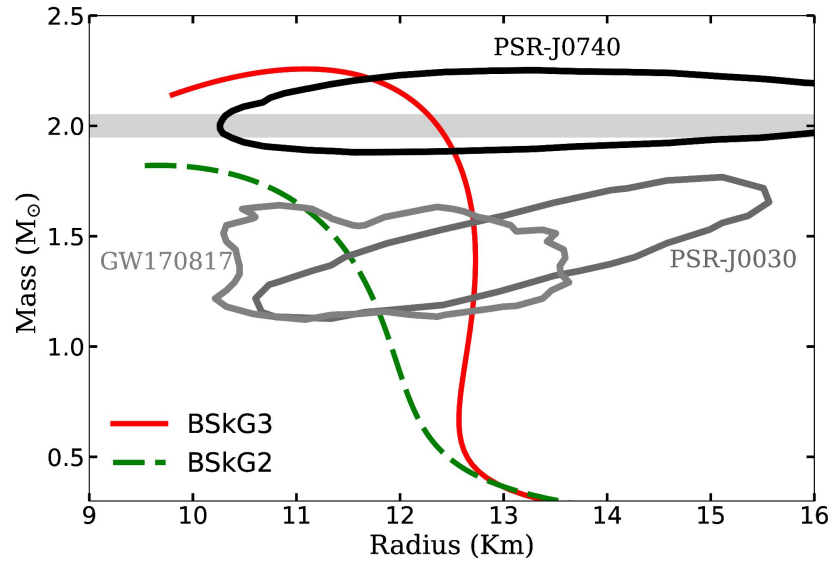
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Fission properties of 45 actinide nuclei

- includes odd-A and odd-odds
- **all** inner barriers exploit triaxiality
- **all** outer barriers exploit
 - octupole deformation
 - triaxial deformation

Neutron stars

G. Grams, W.R. et al., in preparation.

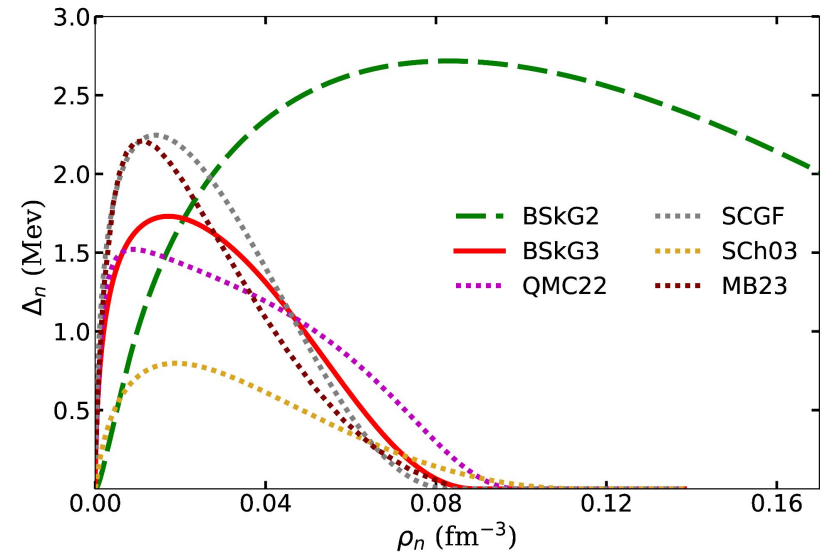
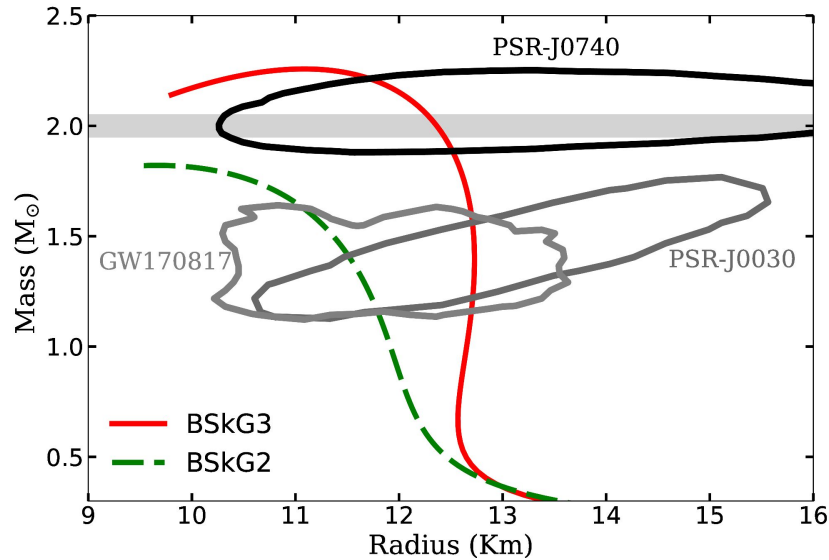


More realistic NS predictions:

- higher maximum mass
 - compatible with NICER
 - compatible with LIGO-VIRGO

Neutron stars

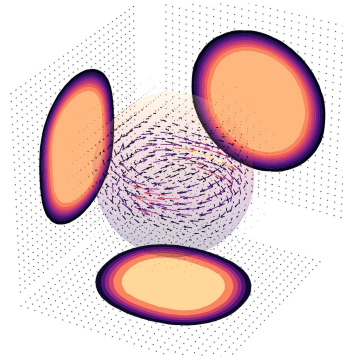
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More realistic NS predictions:

- higher maximum mass
 - compatible with NICER
 - compatible with LIGO-VIRGO
- realistic pairing properties in INM
 - constrained to advanced calculations
- but not at the cost of finite nuclei!

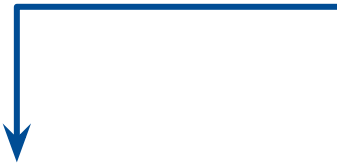
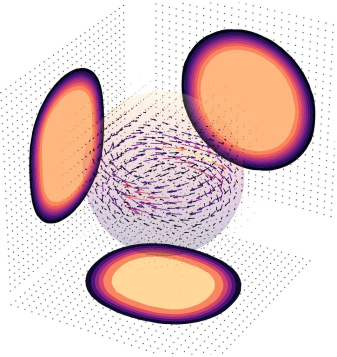
Additional observables



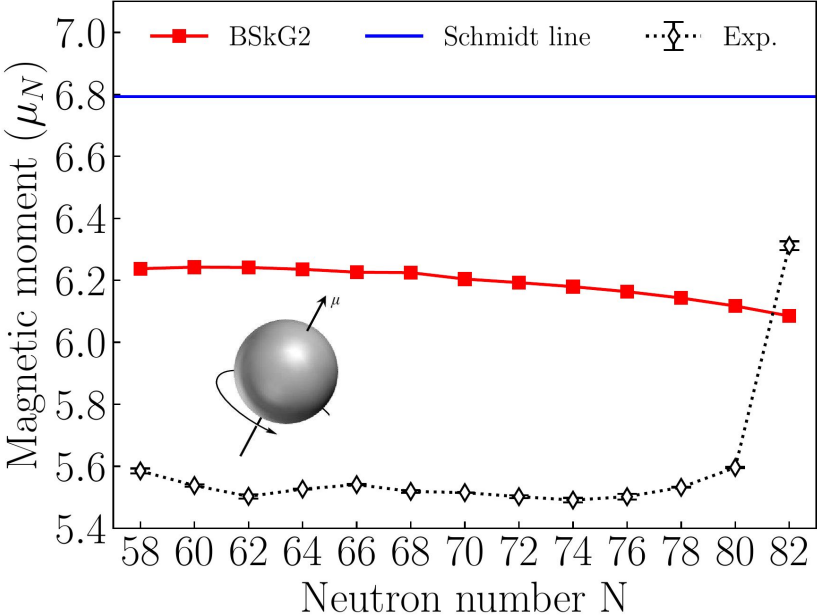
A. R. Vernon et al., *Nature* **607**, **260** (2022),
J. Eberz et al., *NPA* **464**, 9 (1987).
J.Y. Zeng et al. *PRC* **50**, 1388 (1994)

Additional observables

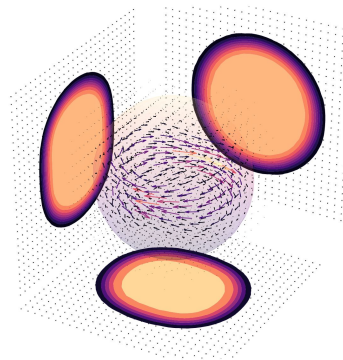
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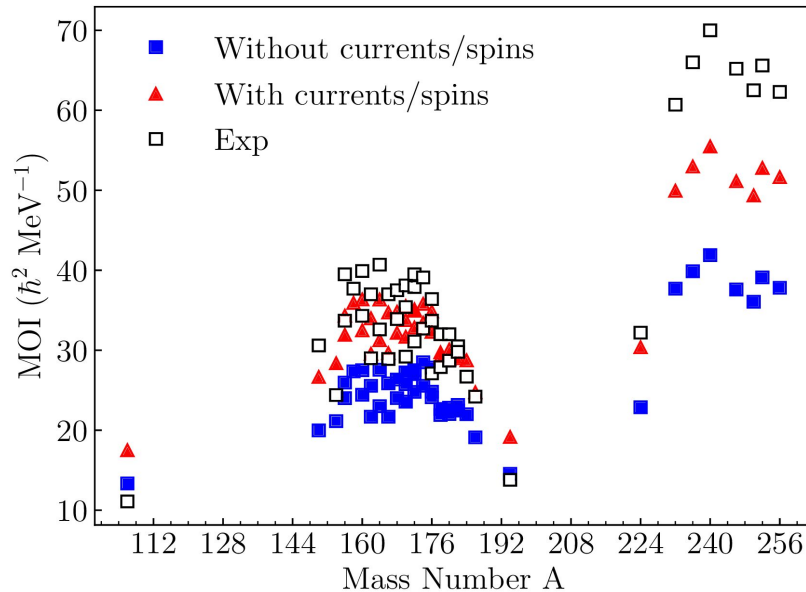
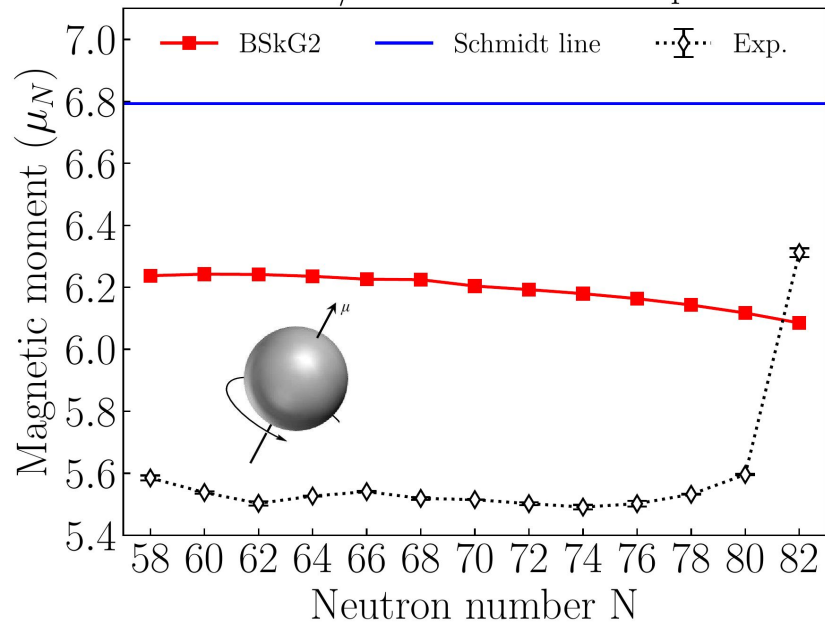
$J^\pi = 9/2^+$ states in In isotopes



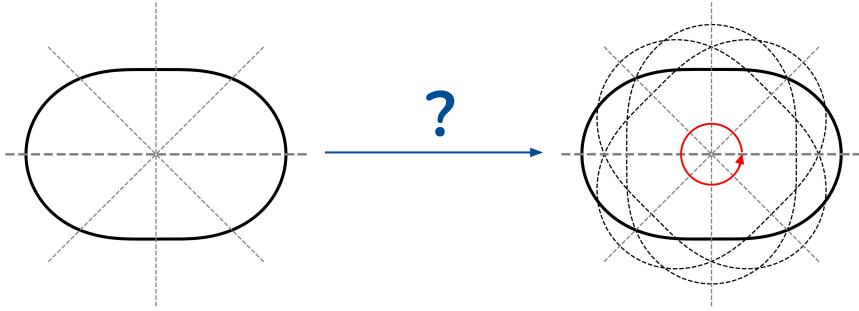
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Challenges: less phenomenology

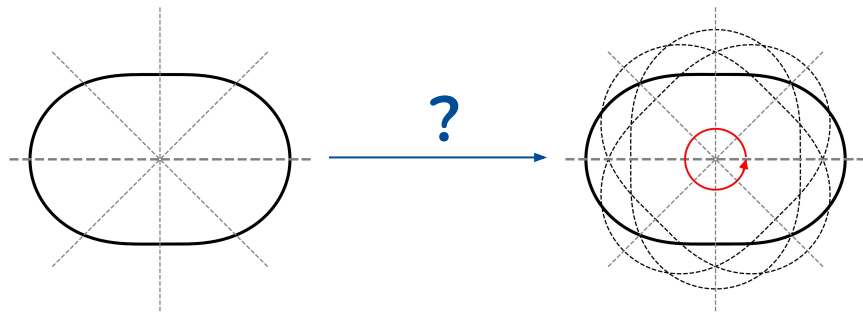


Leave the mean-field picture behind

- techniques exist
- ... but remain extremely costly

Challenges: less phenomenology

$$E \sim \int d^3r \left[C^\rho \rho(\mathbf{r})\rho(\mathbf{r}) + C^\tau \tau(\mathbf{r})\rho(\mathbf{r}) + ?.. \right]$$



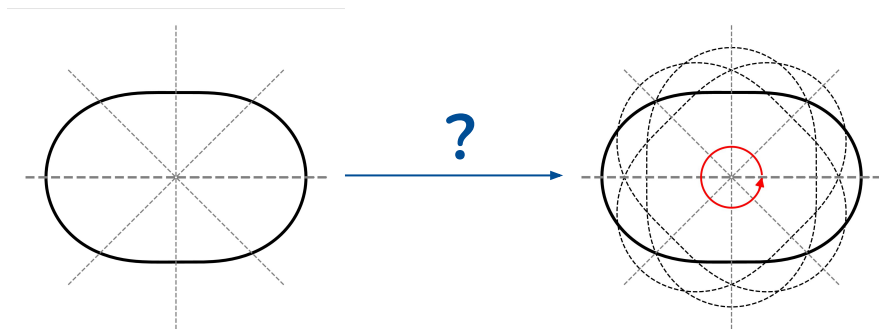
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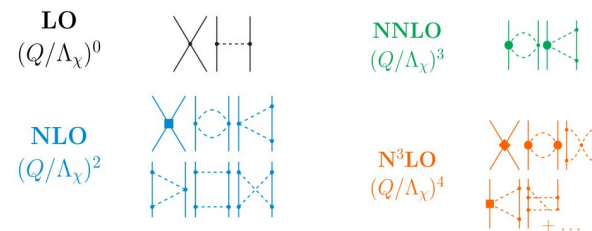
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- ways for **systematic** construction?
- ... perhaps by linking with ab initio?

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Adapted from H. Hergert, Front. Phys. 8:379 (2020).

Leave the mean-field picture behind

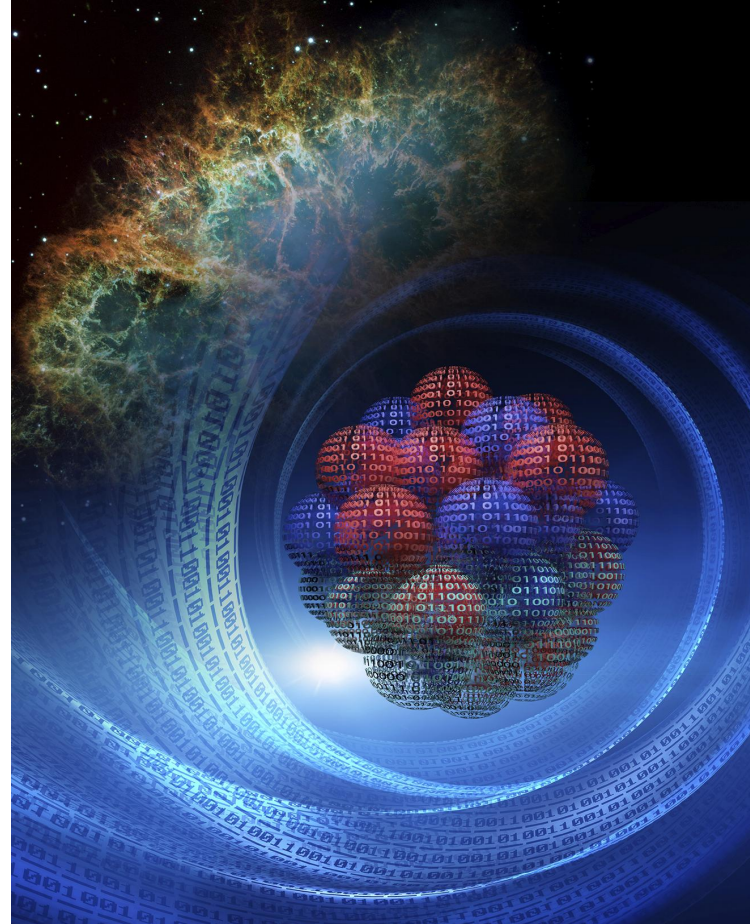
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Conclusion

We build large-scale, microscopic models for (astro) applications.

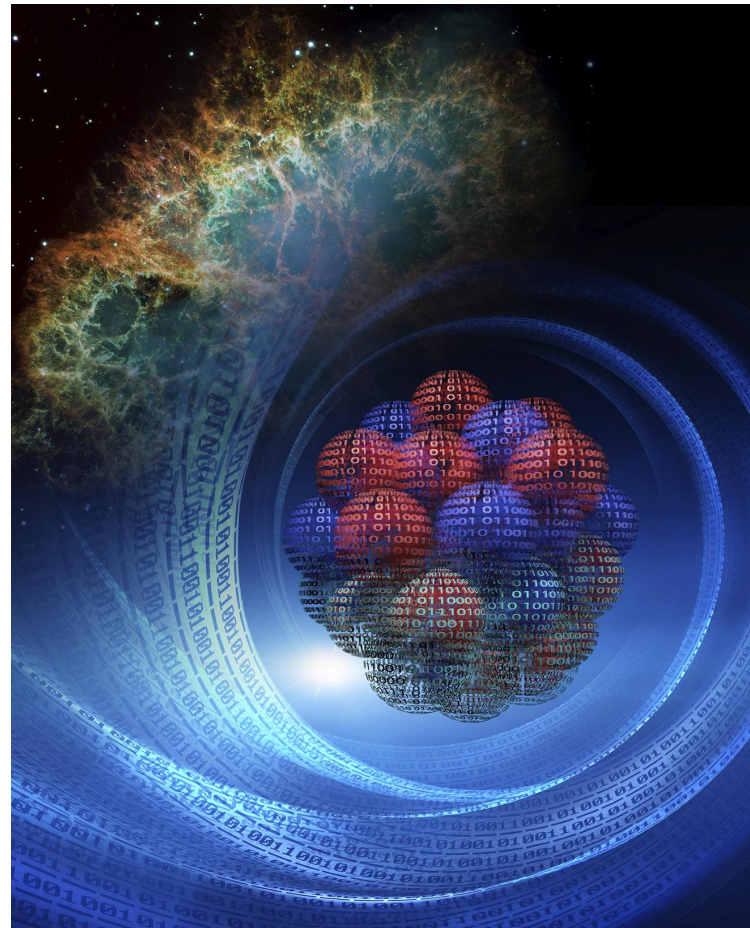


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We build large-scale, microscopic models for (astro) applications.

Large-scale = thousands of nuclei and many observables.

Microscopic = simple wave functions yet complex symmetry breaking.



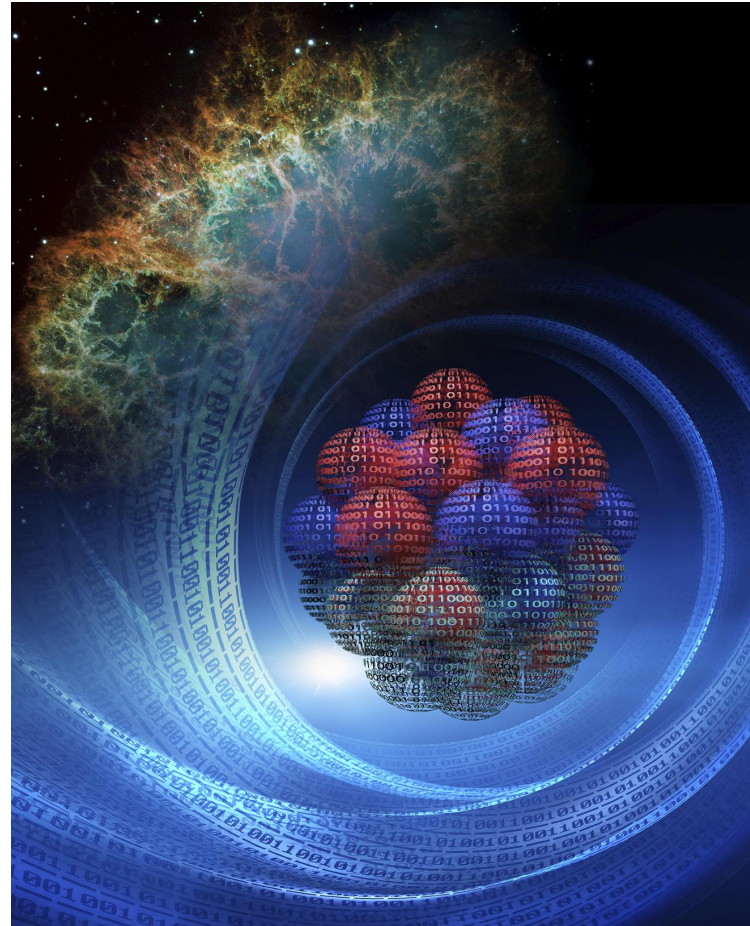
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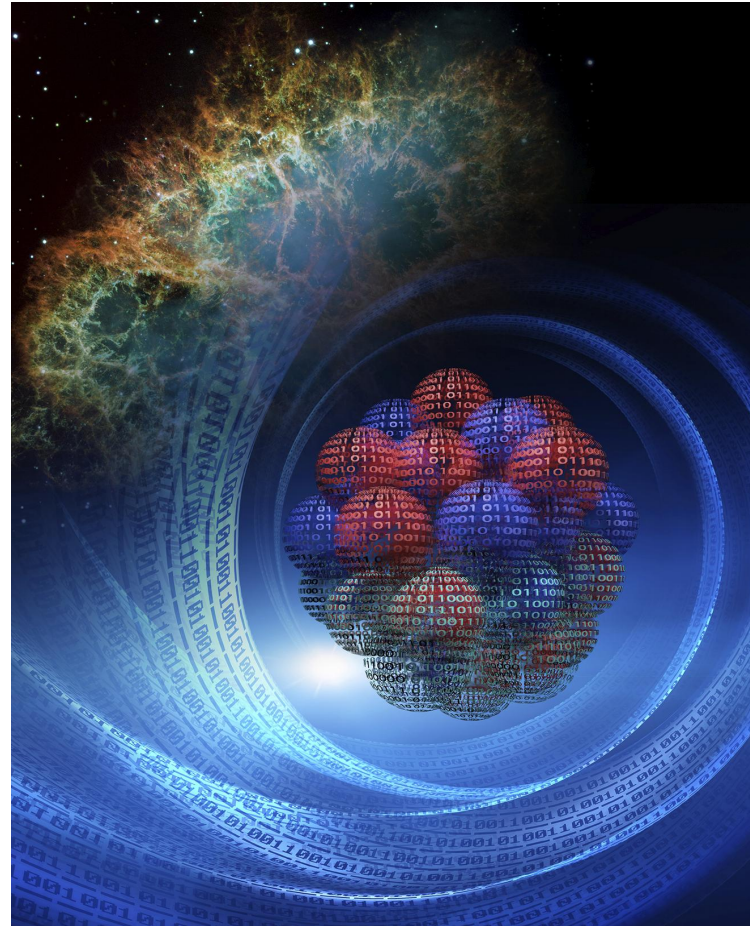
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 - time-reversal breaking
 - octupole deformation
- competitive reproduction of masses and charge radii
- best on the market for **fission** properties
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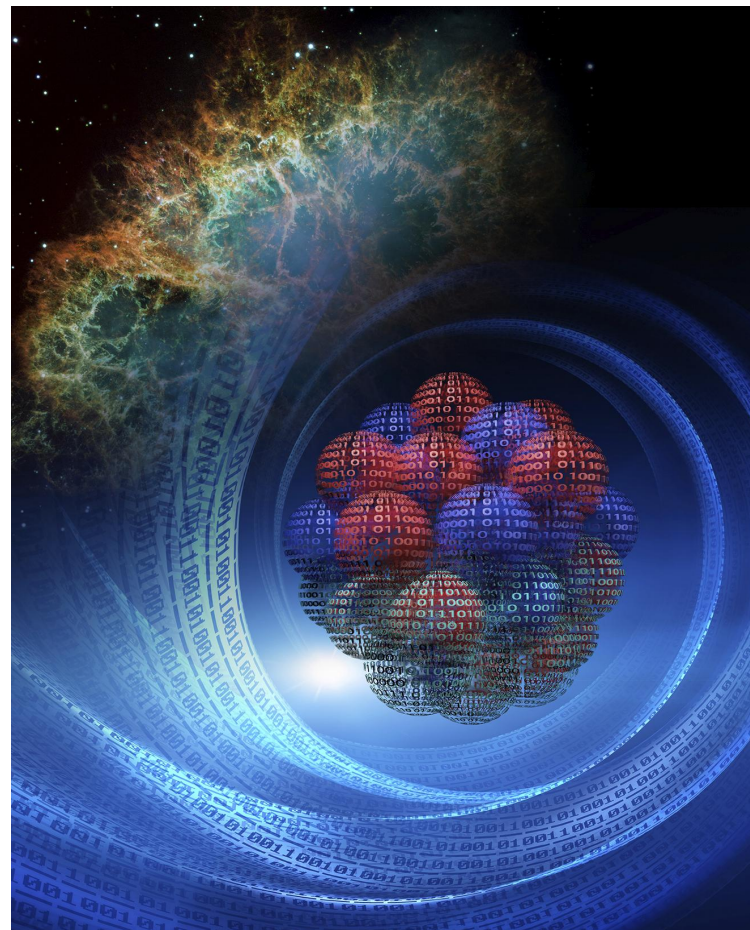
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Coming up from the Brussels group:

- all BSkG3 data
- detailed study of ground state densities
- large-scale **fission** and **level density** calculations
- unified Equation of State for neutron star applications



Thank you for...

..... all the wonderful work!



S. Goriely

G. Grams

N. Chamel

N. Shchepochin



M. Bender

J. Bonnard



G. Scamps



M. Hukkanen

M. Stryjczyk

A. Kankainen



P. Ascher

S. Grévy



E. Verstraelen

T. Cocolios

P. Van Duppen



**UNIVERSITÄT
HEIDELBERG**
ZUKUNFT
SEIT 1386

G. Giacalone



B. Schenke

C. Shen



S. Hilaire

..... the computing time!



..... the funding!



..... your attention!